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# Advancing radioastronomy into the era of High Performance Computing: the case of Imaging

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# RADIOASTRONOMY SOFTWARE

Working at the interface between radioastronomy HPC, my focus is:

- trying to enable the software used for the **processing and analysis of radio data to effectively exploit supercomputing solutions,**
- address the challenge posed by increasingly **larger and complex datasets.**
- **Providing solutions to improve current codes (NOT to replace them) - WSClean**

# MOVING TOWARD HPC

HPC currently means

- Parallel-Distribution
- Parallel-multitasking
- Accelerated computing
- Memory optimization
- Parallel I/O
- Access to Object Storage
- High-performance networking
- Integration with Cloud
- Containers
- .....



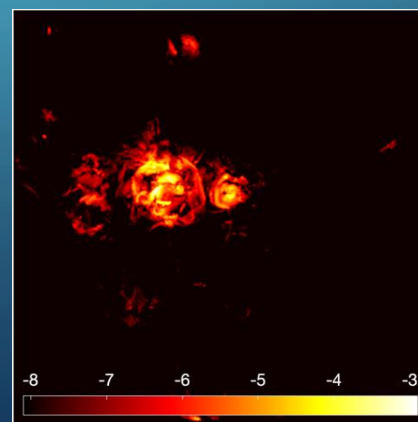
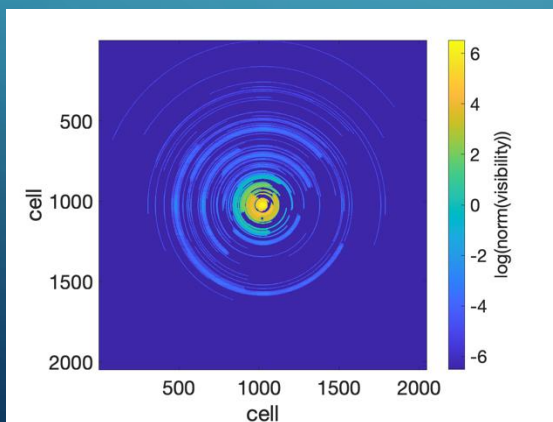
# CASE STUDY: IMAGING

Essentially, we want to invert the following integral:

$$V(u, v, w) = \int \int \frac{I(l, m)}{\sqrt{1 - l^2 - m^2}} \times e^{-2\pi i (ul + vm + w(\sqrt{1 - l^2 - m^2} - 1))} dl dm$$

that **maps** the visibilities  $V$  measured from the interferometer to the sky brightness  $I$ , providing the actual image of the sky.

$(u, v, w)$  are the baselines coordinates, and  $(l, m)$  are the sky coordinates.





# IMAGING: MAIN STEPS

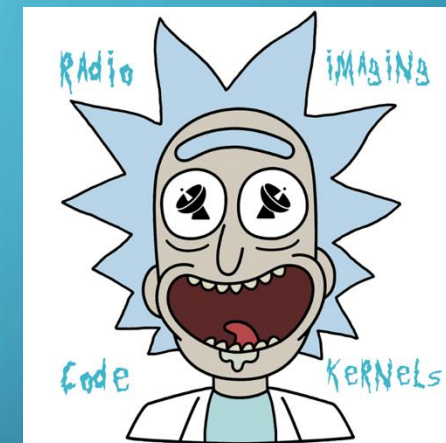
Imaging is a 5D problem: 3 spatial dimensions, frequencies, polarisations.  
It requires essentially 3 **computational demanding** operations:

1. **Discretization** of the problem → map visibilities on a regular mesh (needed for FFT) + weighting + tapering
2. **FFT** transform from Fourier to Real space
3. (**W-correction**, if needed, to correct for Earth curvature)

These operations represent computational challenges that can benefit from High Performance Computing.

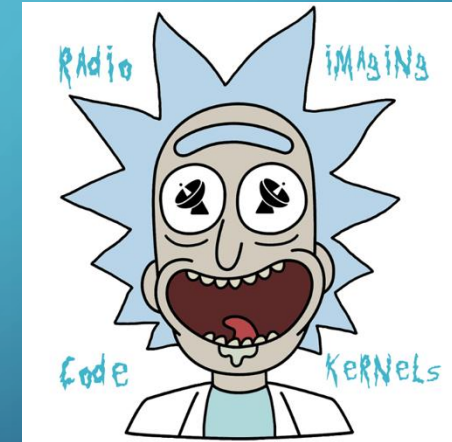
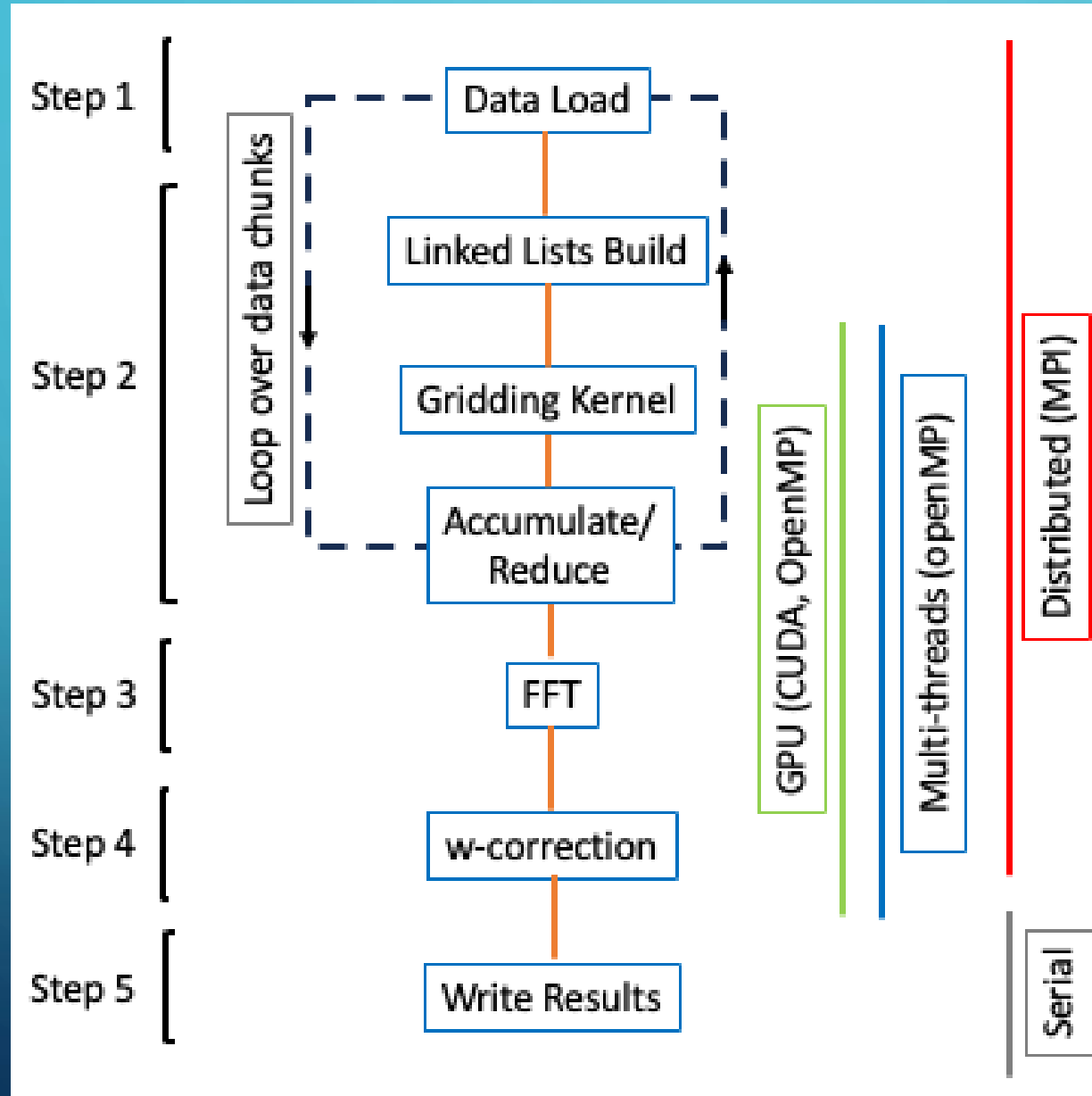
# RICK (Radio Imaging Code Kernels)

- RICK (Radio Imaging Code Kernels) is a code that addresses the **gridding**, **FFT** and **w-correction**, combining parallel and accelerated solutions.
- It is being designed not to substitute radioastronomy codes but to provide specific solutions, portable and fast
- **C, C++, CUDA, HIP** (for AMD GPUs)
- **MPI & OpenMP** parallel, fully working in parallel
- It can run on GPUs (both CUDA and OpenMP for GPU offloading), in particular the FFT using the distributed CUDA library **cuFFTMp**
- An optimized version of the reduce has been developed on both CPU (combining MPI+OpenMP) and GPU (using NCCL or RCCL, for Nvidia and AMD respectively)



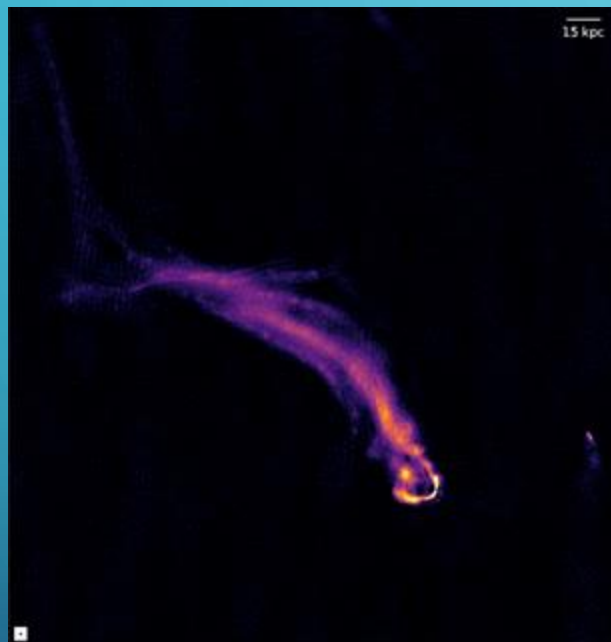
# RICK HPC architecture

The code is publicly available at [https://www.ict.inaf.it/gitlab/claudio.gheller/hpc\\_imaging](https://www.ict.inaf.it/gitlab/claudio.gheller/hpc_imaging)



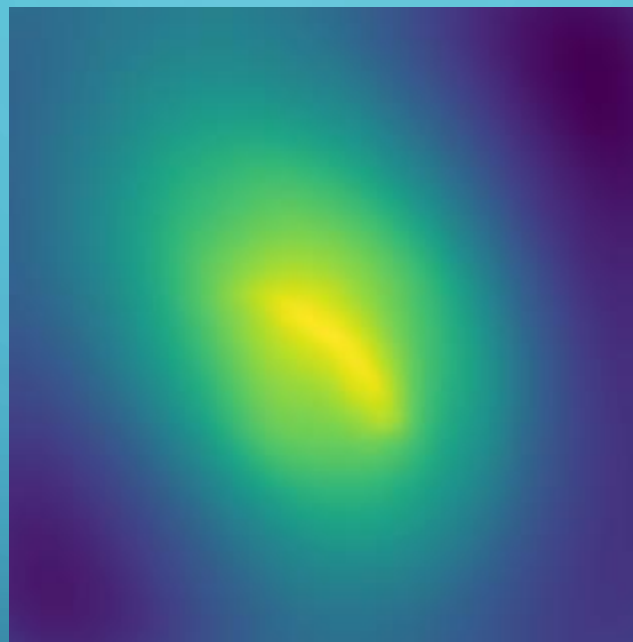
# RICK: DOES IT WORK?

Deconvolved image

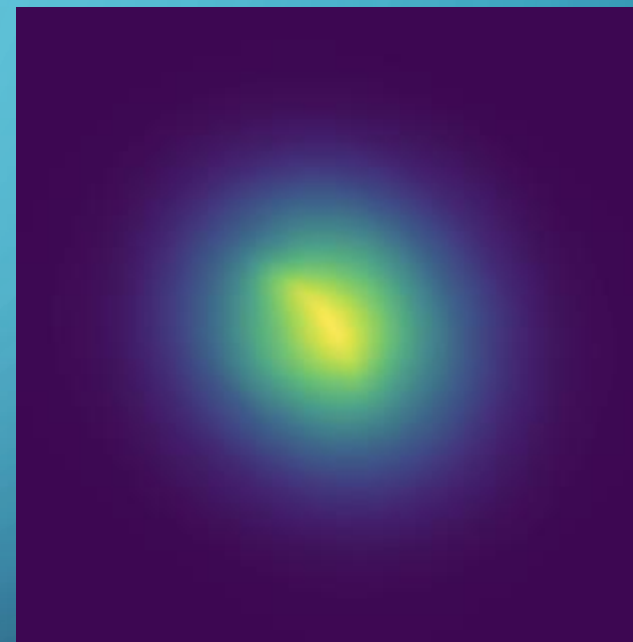


The dataset refers to the "Original TRG", a head-tail radio galaxy in the galaxy cluster Abell 2255

Dirty image, weighting natural, size: 4096x4096x100



WSClean



RICK



# THE COMPUTING SYSTEM: LEONARDO@CINECA

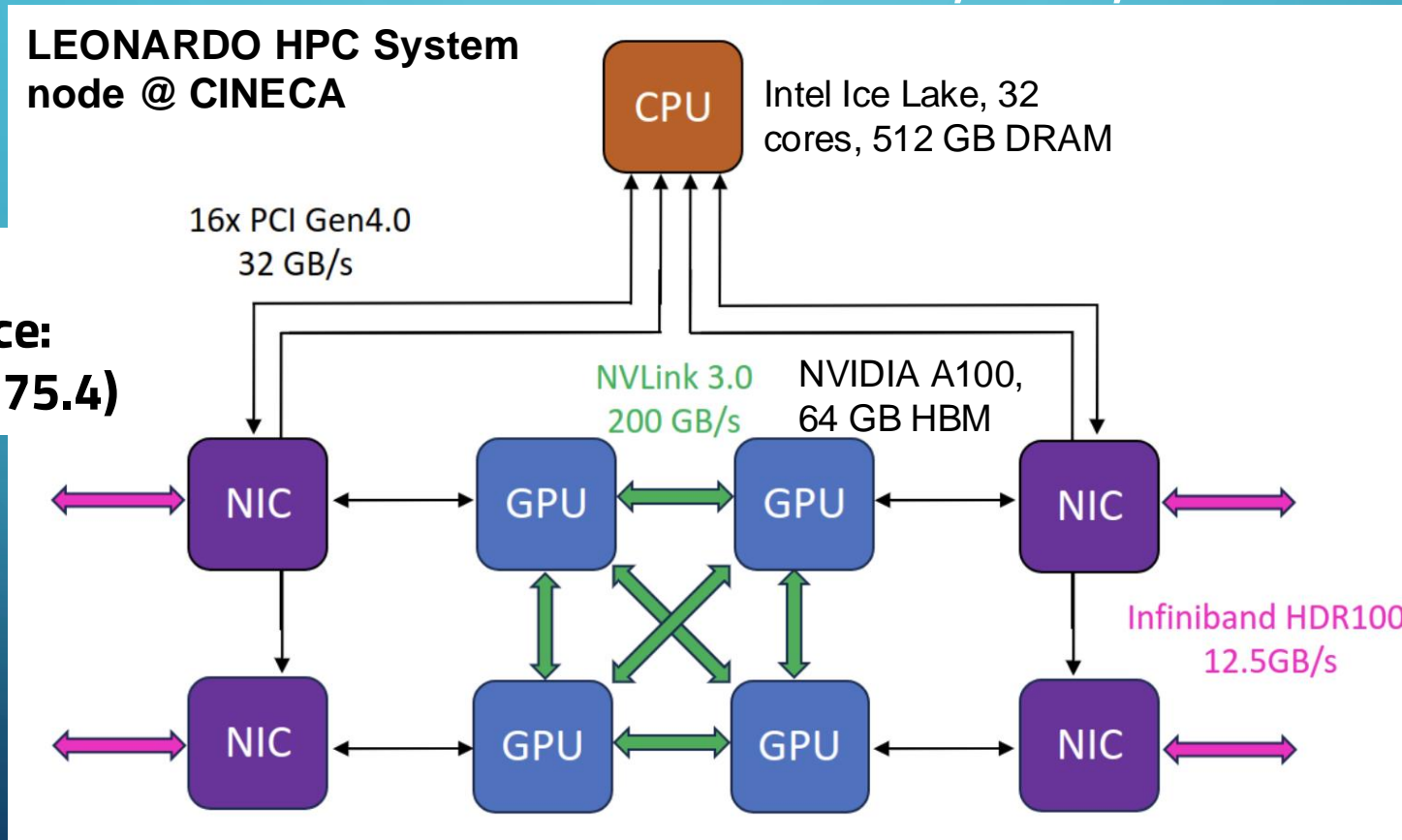


- **4992 Nodes**
- **~300 PFlops**

Showing only a few results. For more, see:

- De Rubeis et al, *Astronomy and Computing*, 2024
- Lacopo et al, submitted
- Gheller et al, *SPIE proc.*, 2024
- Gheller et al., *RASTI*, 2023

**Nominal Performance:**  
**78 TF (2.6+75.4)**



Turisini et al, 2024

# SINGLE NODE TEST

## LOFAR HBA Inner Station:

- 1891 baselines
- 146 MHz, 4 polarizations
- 8 hrs

## Data:

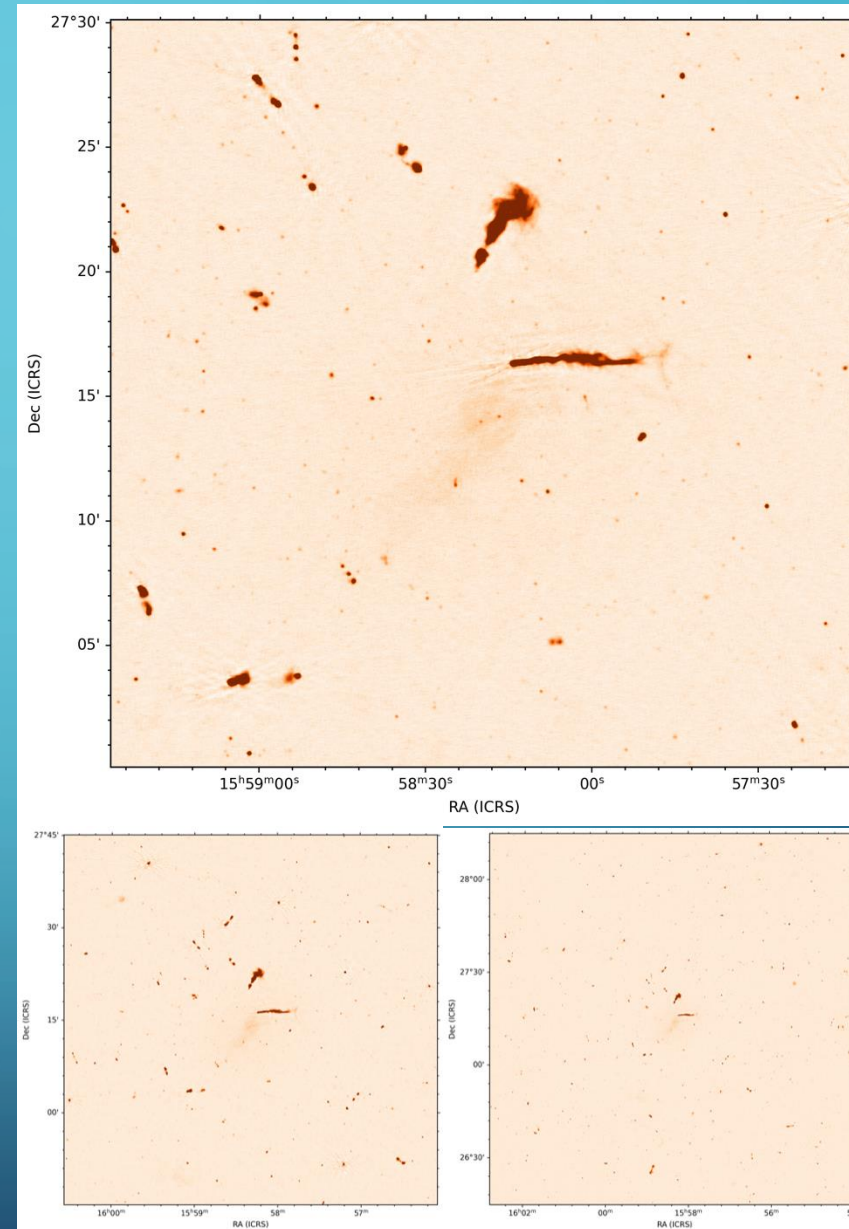
- 543 million visibilities
- 4.5 GB

## Mesh/Image:

- 4096x4096 px, 80 MB
- 4096x4096x16, 430 MB

## Memory Usage

- ~ 10 GB



Courtesy of Luca Bruno



# SINGLE CORE (single node test)

Our BASELINE

Threads	GPUs	Gridding [sec]	FFT [sec]	w-stack [sec]	Comm. [sec]	Total [sec]
1	0	$322.06 \pm 2.95$	$30.01 \pm 0.54$	$26.08 \pm 0.03$	0	$418.07 \pm 3.25$

# MULTICORE (single node test)

OpenMP Multithreading allows to use all the cores of a CPU, accessing all its memory

Threads	GPUs	Gridding [sec]	FFT [sec]	w-stack [sec]	Comm. [sec]	Total [sec]
1	0	322.06± 2.95	30.01±0.54	26.08± 0.03	0	418.07± 3.25
32	0	12.26± 0.14	3.07±0.07	2.18± 0.19	0	48.30± 0.80

29x

10x

12x

9x

# GPU ACCELERATION (single node test)

Comparing FULL CPU with FULL GPU

Threads	GPUs	Gridding [sec]	FFT [sec]	w-stack [sec]	Comm. [sec]	Total [sec]
1	0	322.06± 2.95	30.01±0.54	26.08± 0.03	0	418.07± 3.25
32	0	12.26± 0.14	3.07±0.07	2.18± 0.19	0	48.30± 0.80
1	1	23.29± 0.02	0.70±0.01	0.1429± 0.0002	0	68.98± 0.61
1	2	9.51± 0.07	0.86±0.10	0.0781± 0.0030	1.80±0.01	47.34± 1.05
1	4	4.53± 0.05	0.53±0.01	0.0403± 0.0008	1.51±0.22	23.67± 0.44

2.7x

6x

55x

COMM.

2x



# LOFAR-VLBI USE CASE

LOFAR-VLBI represent a good benchmark to SKA in terms of data volume, with ~15 TB of end-products for a single, 8hrs, observation.

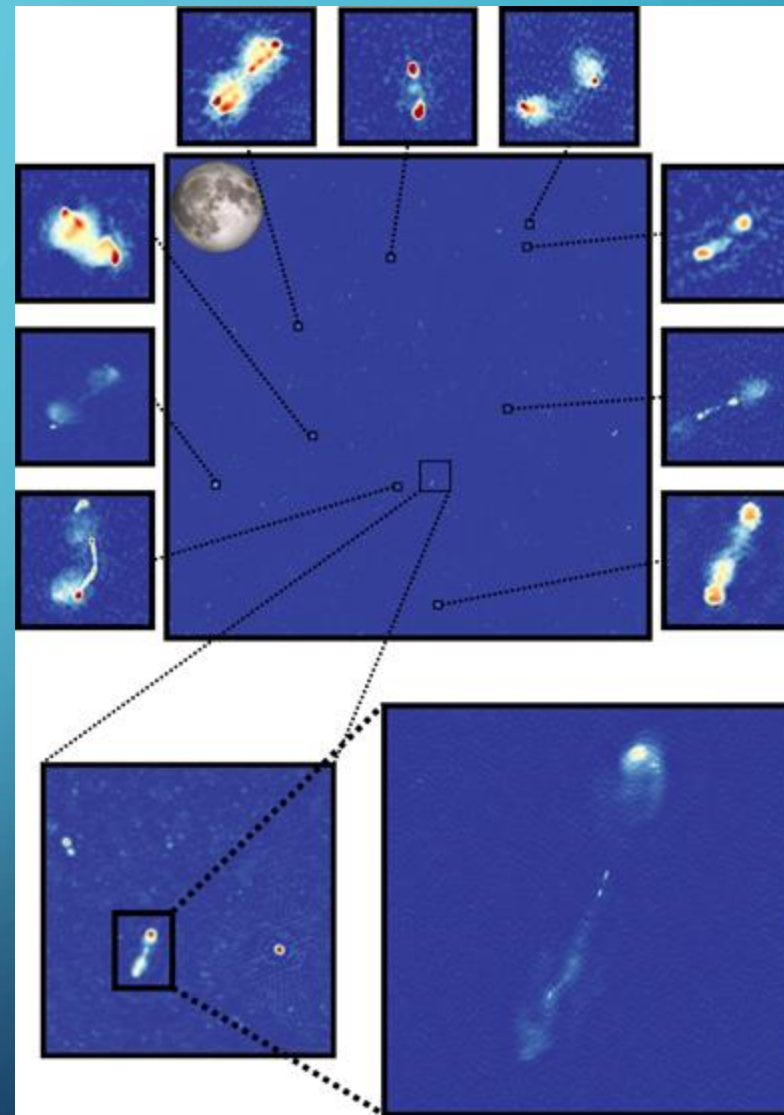
An interesting and challenging task is the imaging of a large FoV (~2.5deg x 2.5deg) with sub-arcsecond resolution using LOFAR-VLBI.

For our LARGE tests, we have:

Input data ~ 533 GB

Output Image = 65536 x 65536 x 32 pixels ~ 4 TB

This is not something a single CPU/GPU can face!

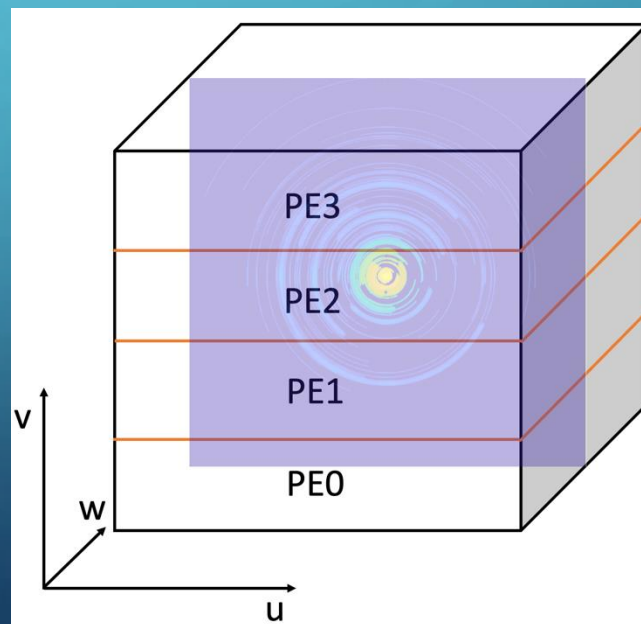
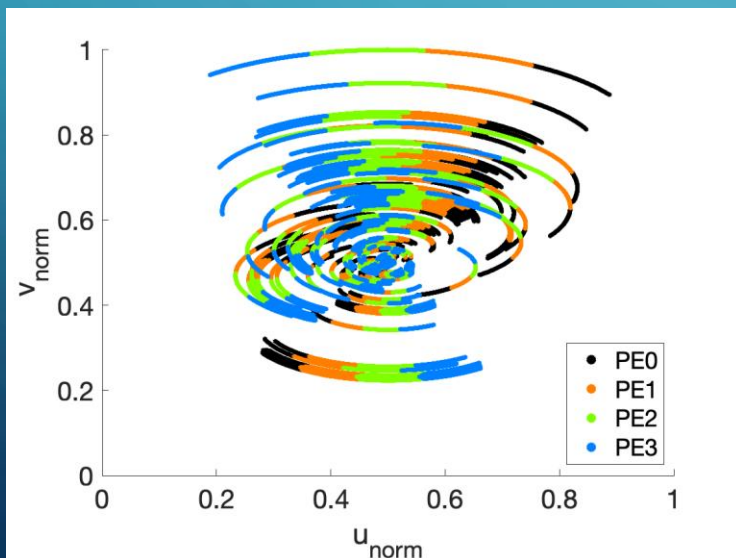


# OVERCOMING THE MEMORY WALL

Parallel computing allows to use multiple processors distributing data among their memory:

- Visibilities (and work) are evenly distributed among processing units
- The mesh is split among processing units. The full mesh is never stored in a single memory

→ Problems of “any” size can be supported



Main issue:

visibilities are distributed across memories unrelated to mesh slabs

→ Lots of communication required

# LEONARDO@CINECA LARGE TESTS



	Nodes	MPI tasks (threads per task)	GPUs	Gridding (s)	Reduce (s)	FFT (s)	w-correction (s)	Total (s)
CPU tests	32	1024 (1)	0	4.5	9631.4	160.6	7.2	10246.0
	64	2048 (1)	0	1.9	9598.2	107.1	3.5	10153.5
	128	4096 (1)	0	1.1	9715.8	98.4	1.7	10266.5
GPU tests	32	128 (8)	128	2.6	54.8	4.2	0.3	67.4
	64	256 (8)	256	2.4	59.4	2.8	0.2	69.4
	128	512 (8)	512	2.7	72.6	2.7	0.1	83.4

- The right balance of computing power and scaling must be determined to be efficient
- COMMUNICATION BECOMES **THE** BOTTLENECK
- GPUs help reducing the communication bottleneck

# WORK IN PROGRESS & CONCLUSIONS

- Accurate Evaluation of the Energy impact of RICK
- Modularization toward library to be used from WSClean
- Optimization of I/O through parallelism
- Integration in complex pipelines (Streamflow - <https://streamflow.di.unito.it/>)

- HPC can help in strongly reducing the time to solution of radioastronomy data processing of **order of magnitudes**.
- **MPI Parallelism is the only** alleviate I/O bottlenecks
- Computing nodes must be algorithmic components. **scalability**

THANKS FOR YOUR  
ATTENTION